

Overview of the Argonne Data Science Program & Science Highlight

Elise Jennings

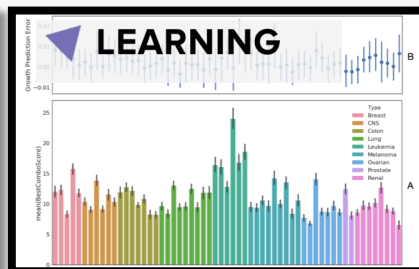
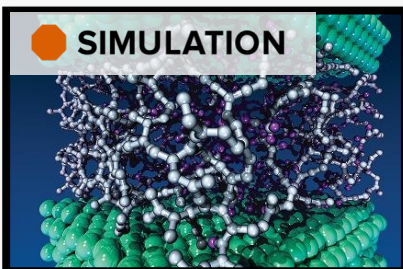
ejennings@anl.gov

Argonne Leadership Computing Facility



The Argonne Leadership Computing Facility provides world-class computing resources to the scientific community.

- Users pursue scientific challenges
- Resources fully dedicated to open science
- In-house experts to help maximize results



ALCF offers different pipelines based on your computational readiness. Apply to the allocation program that fits your needs.

<https://www.alcf.anl.gov>



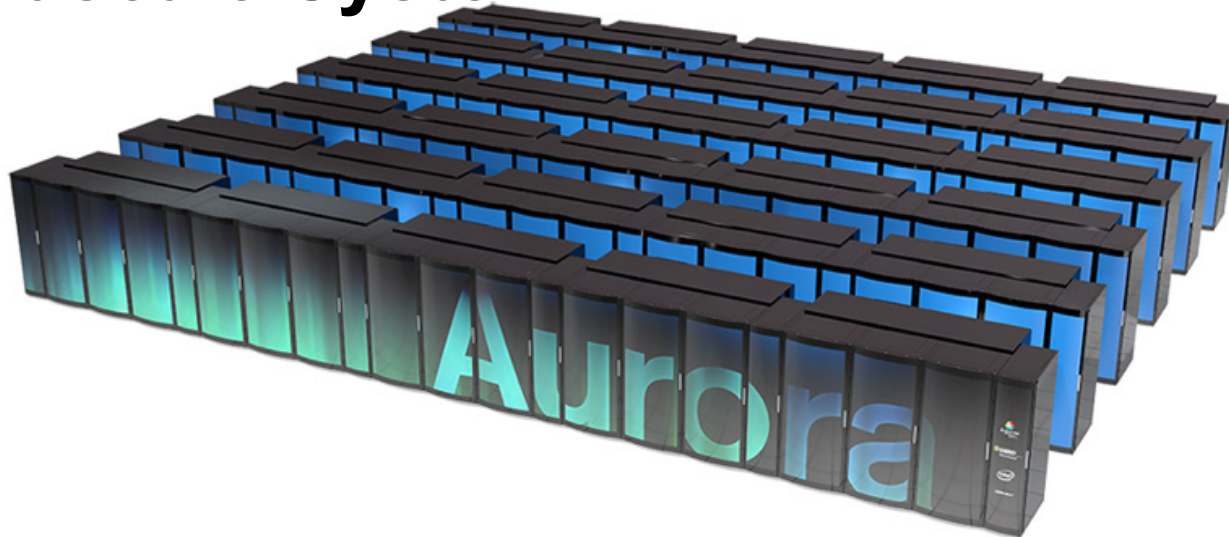
Theta Intel/Cray

4,392 nodes
281,088 cores
69 TiB MCDRAM
824 TiB DDR4
549 TB SSD
Peak flop rate: 11.69 PF

Cooley

NVIDIA system
126 nodes
126 K80 GPUs
Peak flop rate: 293 TF

Aurora 2021 (A21) The first US Exascale System



Architecture supports three ways of computing

- Large-scale Simulation (PDEs, traditional HPC)
- Data Intensive Applications (scalable science pipelines)
- Deep Learning and Emerging Science AI (training and inferencing)

Argonne Leadership Computing Facility





Tom Uram



Venkat Vishwanath



Taylor Childers



Murat Keceli



Elise Jennings



Alvaro Vazquez Mayagoitia



Adrian Pope



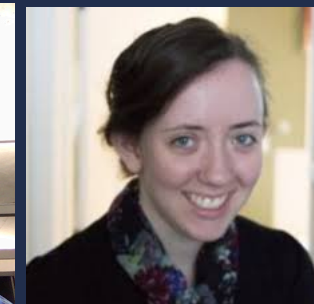
Misha Salim



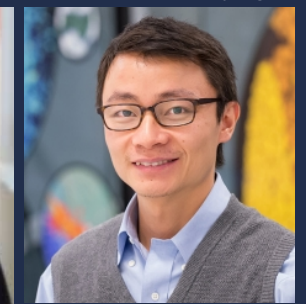
Prasanna Balaprakash



Antonio Villarreal



Bethany Lusch



Huihuo Zheng

datascience@alcf.anl.gov



Himanshu Sharma



Xiao-Yong Jin



Ganesh Sivaraman



Corey Adams

ALCF Data Science Program (ADSP)

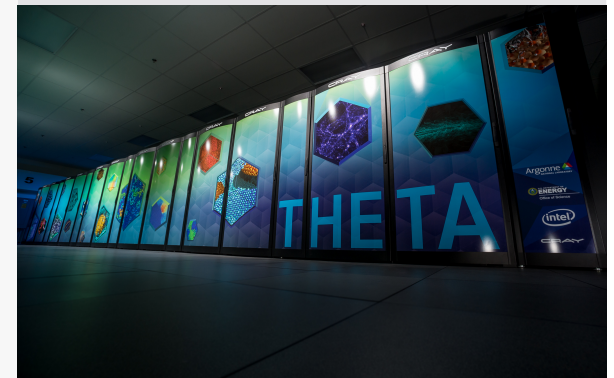
The ADSP program started in 2016 and is now in its 4th year.

ADSP's goal is to support "Big Data" science that require the scale and performance of leadership computing.

Successful projects have

- high potential impact
- data scale readiness
- diversity of science domains and algorithms
- can fully exploit the architectural features of Theta

Two-year proposal period. Yearly call for proposals.



ALCF Data Science Program (ADSP)

Two main targets for development

Science applications

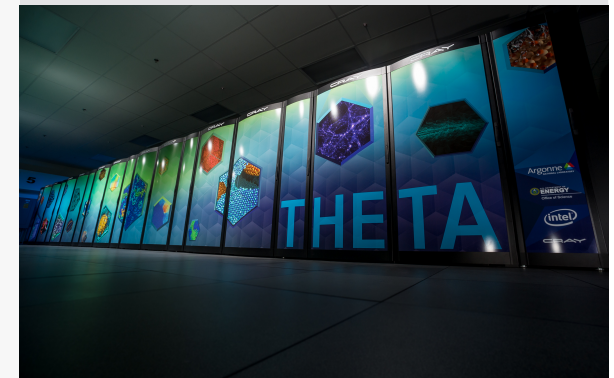
Tools

To date the majority of proposals have been science applications.

Tool development and support is becoming a major requirement.

ADSP projects

- span a diverse set of science domains (Materials, Imaging, Neuroscience, Engineering, Combustion/CFD, Cosmology).
- involve large science collaborations (multiple APS, LSST, DESC, LIGO, DES, ATLAS) and smaller research groups developing ML at scale.
- have used nearly 300M core hours on Theta (26% as capability runs)



Emerging trends

We now have a mix of applications at scale

HPC simulations, Big Data analytics and ML

Deep integration of HPC simulations and Machine/Deep Learning

Augment training data, provide supervised labels for training

ML model can be embedded into the simulation

Speed/accuracy trade off in replacing first principal model with ML

Big drive for

Scientific techniques (Uncertainty Quantification, reproducibility etc.)

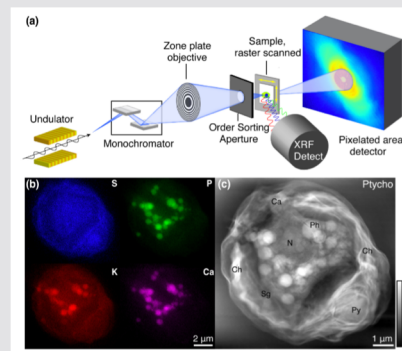
Image processing, in situ analysis and visualization

Complex and interactive workflows with performance capabilities

Smart configuration space sampling

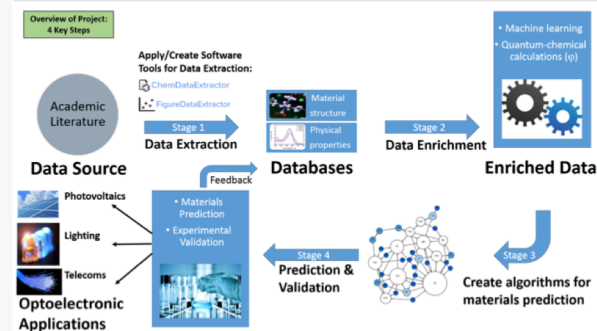
Tools to couple simulations with analysis and ML

X-ray microscopy of extended 3D objects: scaling towards the future PIs: Chris Jacobsen, Wild, Nashed (ANL, Northwestern)



Molecular Engineering of Solar-Powered Windows

PI: Cole, Cambridge University, ANL



Keeping up with the pace of Machine Learning is challenging

The pace of Machine Learning is very different to traditional HPC.

ML/DL software

Updates occur every few weeks (TensorFlow, Keras, PyTorch, Horovod, etc.)

Stack must enable performance libraries (Intel MKL, MKL-DNN, LibXSMM)

Must work seamlessly with simulation and data frameworks

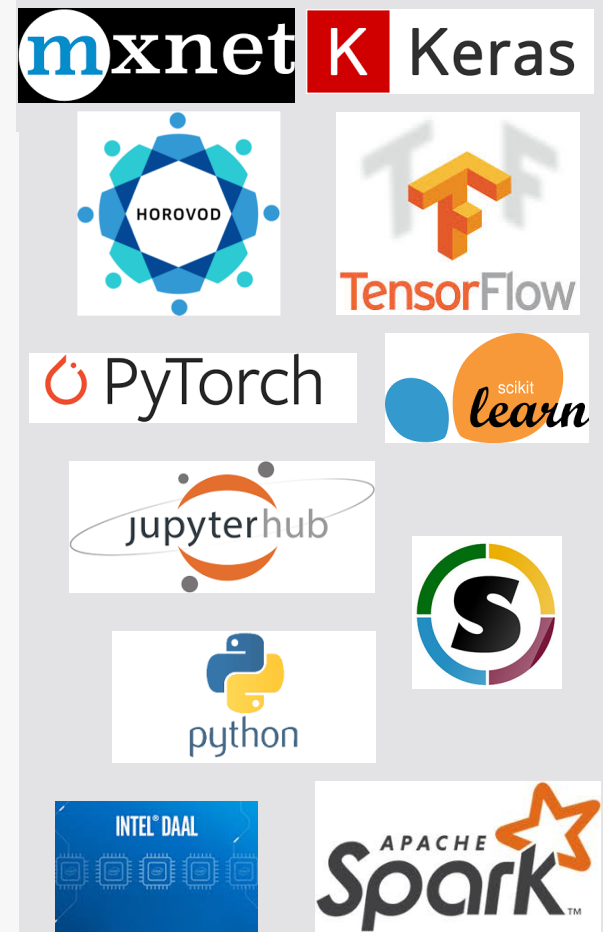
Development does not always prioritize backwards compatibility

Keeping up requires

Dedicated team members to track and update software regularly

Containers which can provide portable customized software stack

Regular training/workshops to update the scientific community



Machine Learning, Deep Learning & Workflow software

ML/DL:

TensorFlow, Keras, Neon, MXNet, Caffe2, Theano, CNTK, PyTorch, Sci-kit Learn, Graph Analytics (Cray Graph Engine), Horovod

With performance libraries e.g. Intel MKL, MKL-DNN, LibXSMM enabled

Intel optimized Tensorflow

Conda package on Theta

Intel Distribution for Python's optimized numpy

Uncertainty Quantification: Edward, Tensorflow Probability

Containers: Singularity

Data Analysis: MongoDB, Apache Spark, R, Python, Jupyter Hub



High impact Data Science software

Balsam workflow manager

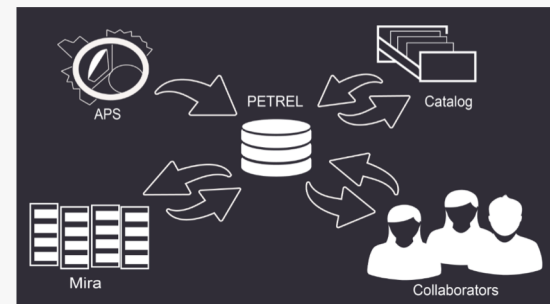
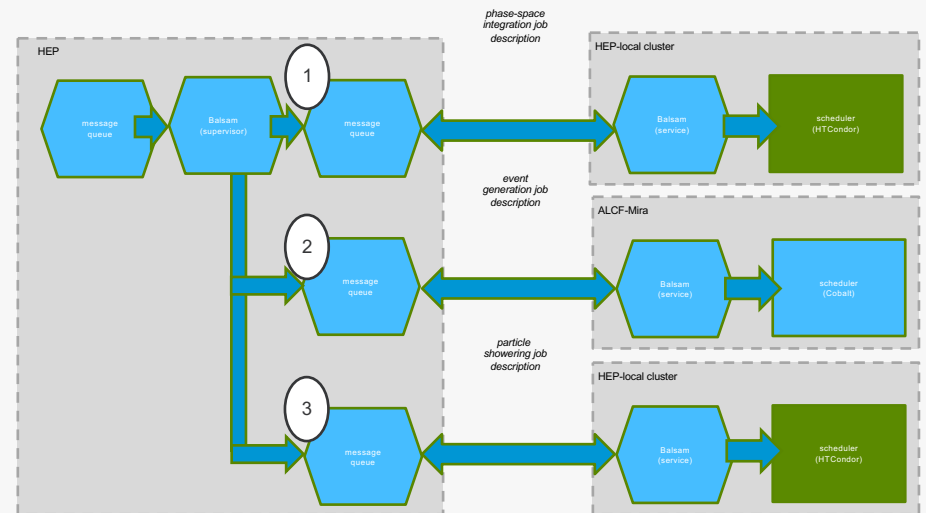
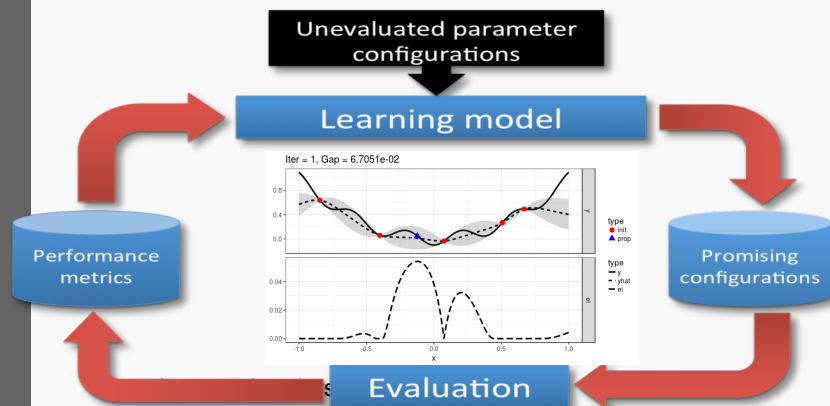
ATLAS experiment used Balsam to run ~100's million compute hours of jobs across ALCF and NERSC systems. Balsam is used by ADSP, ESP, ALCC and ECP applications.

<https://www.alcf.anl.gov/balsam>

Deep Hyper

ALCF is currently conducting hyperparameter optimization for DL on thousands of nodes of Theta

<https://github.com/deephyper/deephyper>



PETREL
Data Management and Sharing Pilot

Petrel

Petrel leverages storage and infrastructure provided by ALCF and Globus Transfer and Sharing services. 100TB allocation per project.

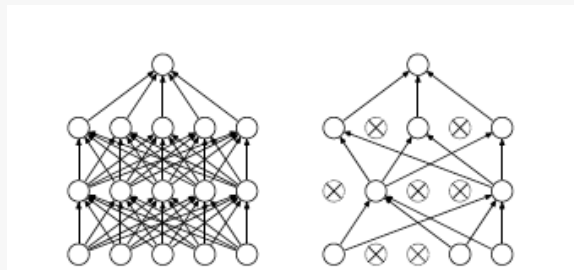
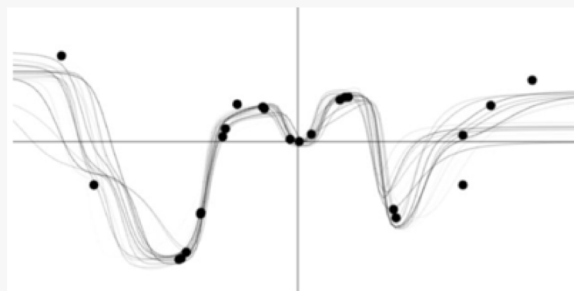
<http://petrel.alcf.anl.gov>

Scientific Machine Learning

Uncertainty Quantification

Software and method development for

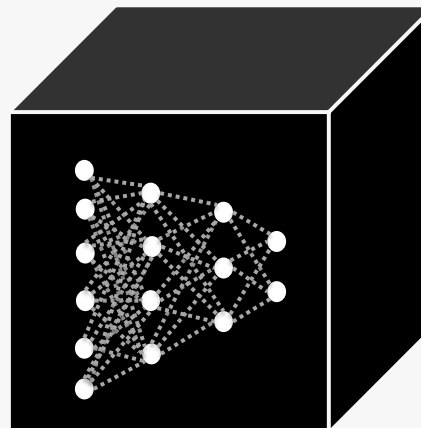
- Bayesian Neural Networks at scale
- Drop out methods




Interpretability

Data Science team provide support on

- visualization of neural network
- information theoretic model of neural networks
- Domain knowledge to develop specific network, loss function



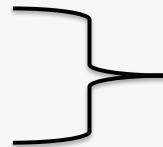


Science highlight: Deep Learning in astrophysics

Deep Learning in Cosmology/Astronomy

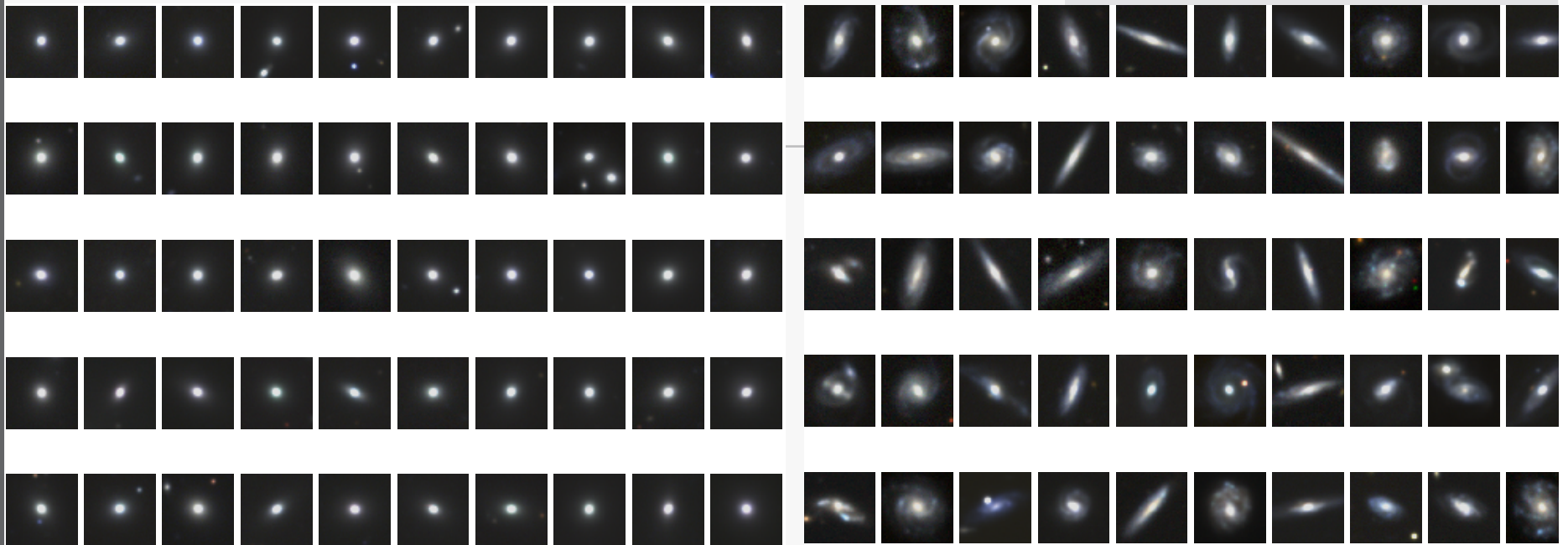
Applications:

- Mock catalogue creation
- Augment N-body simulations
- Detect gravitational lenses
- Detect transients in real time
- Detect gravitational waves in real time
- Classify galaxy images
- Parameter estimation from time series data



**Multimessenger Astrophysics through
the NCSA-Argonne Collaboration**
PIs: Huerta, Zhao, Haas, Saxton (NCSA)

Deep Transfer Learning to classify galaxy images



Deep Transfer Learning to classify galaxy images

Network:

Xception + a few custom defined fully connected layers

Weights:

pretrained weights with the ImageNet dataset

Data:

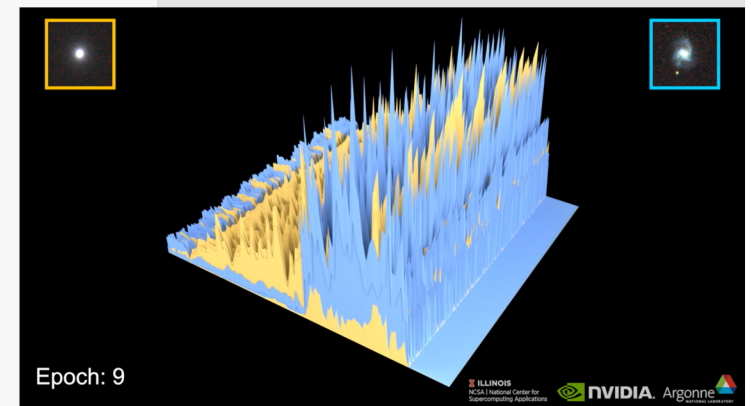
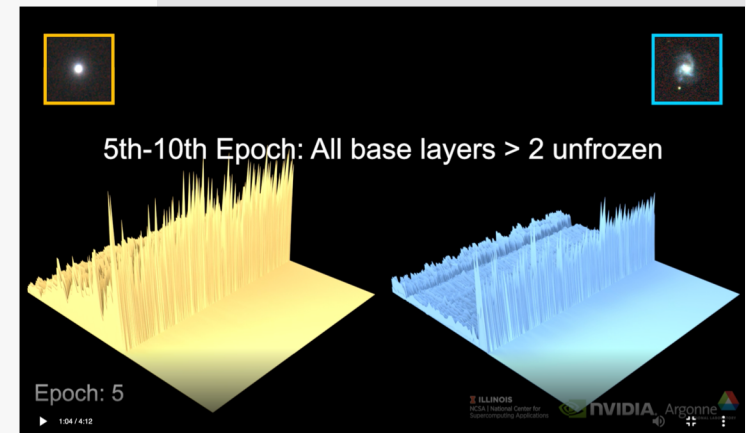
resized all the galaxy images 299×299 pixels

Method:

Progressively unfreeze earlier layers of the whole network

Fine tune their weights for a few epochs of training

Retain earlier layers of a trained network: versatile filters for features like lines and edges



Using the network as a feature extractor

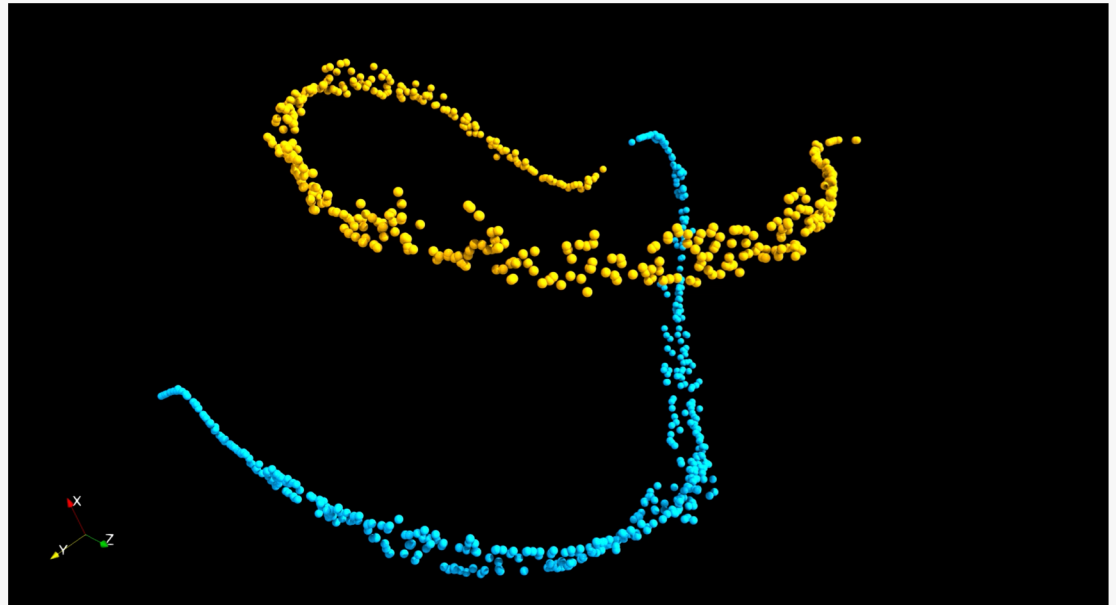
Network output:

output activation values from
second-to-last layer

3D representation

t-Distributed Stochastic Neighbor
Embedding (t-SNE)

Addresses common problem of large
unlabeled datasets



SDSS spirals

SDSS ellipticals

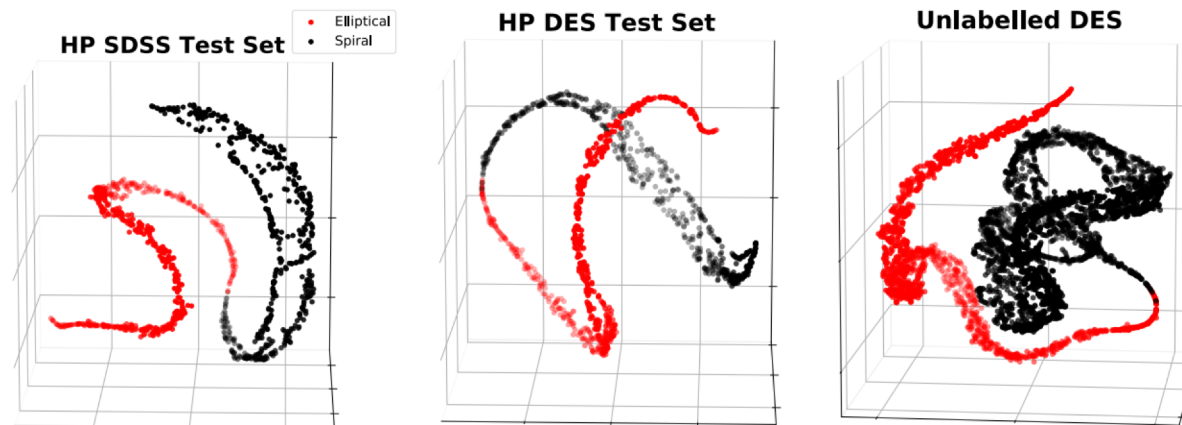


FIG. 4: t-SNE visualization of the clustering of HP SDSS and DES test sets, and unlabelled DES test.

ADSP mmaADSP
Khan et al 2019

Gravitational Wave detection and parameter estimation

A three-detector observation of gravitational waves from a binary black hole coalescence

simulation and scientific visualization by

Gabrielle Allen, Roland Haas, Eliu Huerta, Edward Seidel
Gravity Group
National Center for Supercomputing Applications
University of Illinois at Urbana-Champaign

Novel data-parallel deep learning fusing HPC and AI for MultiMessenger Astrophysics (MMA).

Huge potential for scientific discovery

- Convergence of all-sky GW observations (LIGO) with deep, high-cadence electromagnetic observations (LSST)
- Novel visualization of Neural Networks

Deep Learning for Multi-Messenger Astrophysics. A Gateway for Discovery in the Big Data Era, Huerta et al., Nature Review Physics

Bayesian Neural Networks at scale

- Deep Learning at scale for parameter estimation of Binary Black Hole (BH) mergers
(spins are aligned or anti-aligned, evolve on quasi-circular orbits)
- L2loss re-defined to be negative Evidence Lower Bound (ELBO) loss.

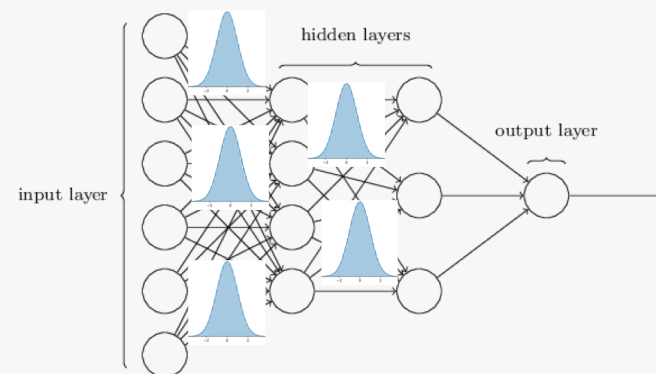
ELBO loss = expected negative log-likelihood +
Kullback-Leibler divergence

- Variational Inference method

Posterior distribution parameter fit by network
Prior distributions for all network parameters

$$w_i^1 \sim \mathcal{N}(0, \epsilon^1)$$

$$w_i^2 \sim \mathcal{N}(0, \epsilon^2)$$



Bayesian Neural Networks at scale

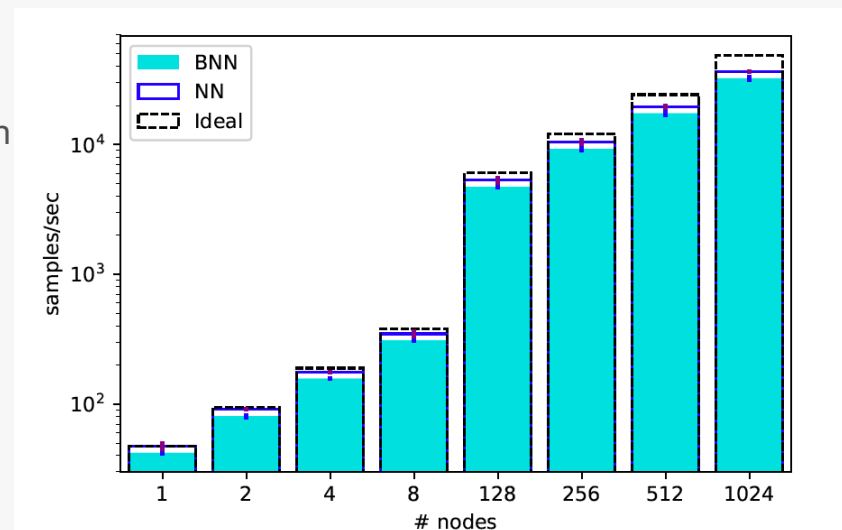
Inference is very costly

- Negative Log Likelihood is approximated via Monte Carlo
- Validation, test steps only carried out when needed

Scaling on Theta with Horovod.

- Horovod and Tensorflow timeline show increased time in AllReduce due to increased parameter set.

Tensorflow Probability on Theta



Thank you !

datascience@alcf.anl.gov

